

Solar and Stellar Flares - nanoflares to superflares -

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thanks to H. Maehara, T. Shibayama et al.

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- Solar flares
 - Recent observations
 - Unified model
- Stellar flares
- Superflares on Solar type stars

Solar flares

Solar flare

Discovered in 19c

Explosive energy release

That occur near sunspot

**magnetic energy is the
source of energy**

Size $\sim 10^9 - 10^{10}$ cm

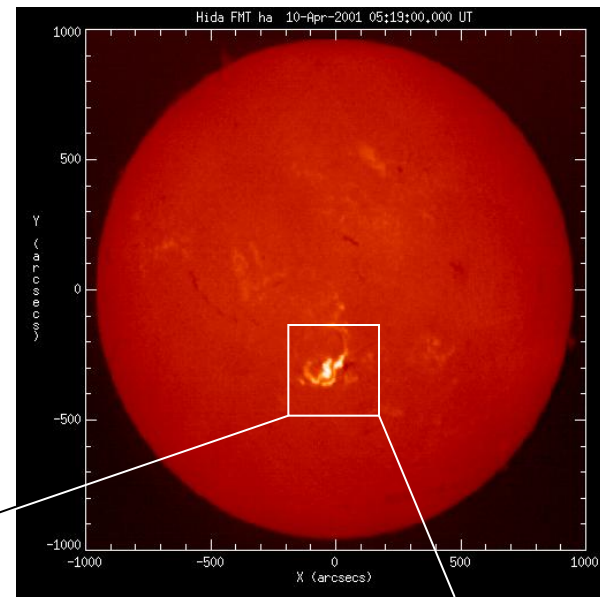
Time scale $\sim 1\text{min} - 1\text{hour}$

Total energy \sim

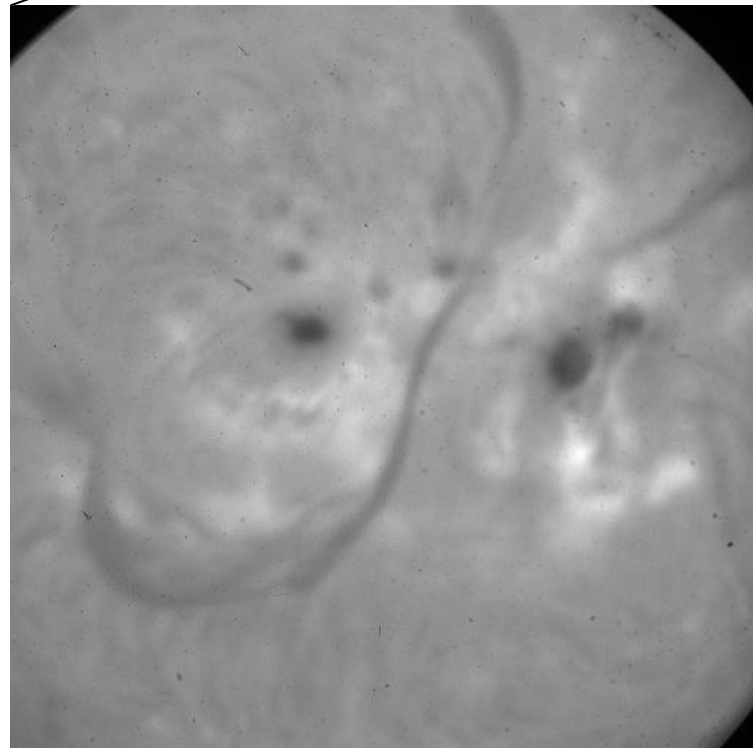
$10^{29} - 10^{32}$ erg

Mechanism has been
puzzling since 19c until
recently

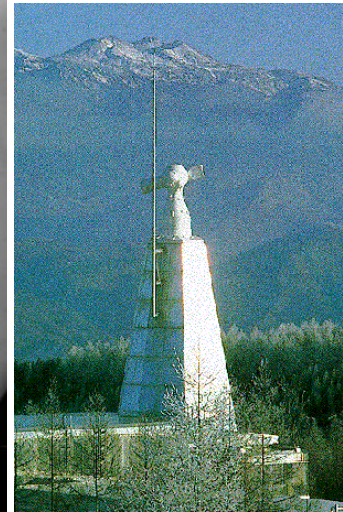
Hida Obs/Kyoto U.



H α



2001.04.10 04:29:26[UT]





Solar corona observed in soft X-rays (Yohkoh)

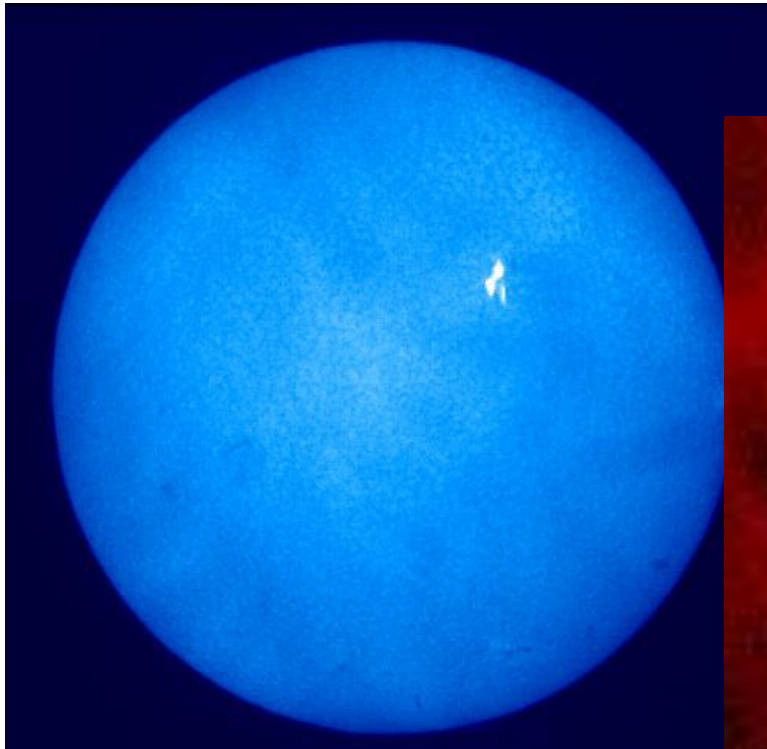
Soft X-ray
telescope
(1keV)

Coronal
plasma
2MK-10MK



Simultaneous H α and X-ray view of a flare

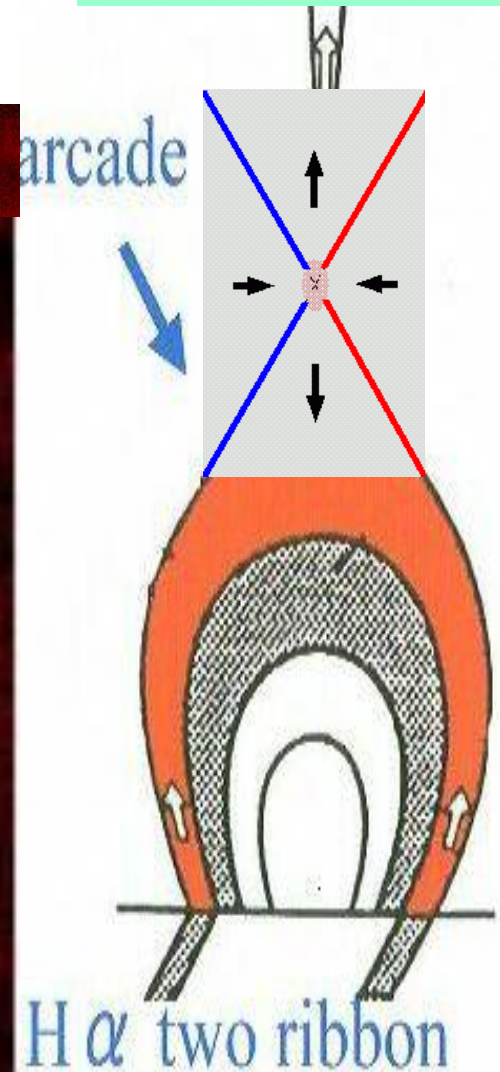
H α



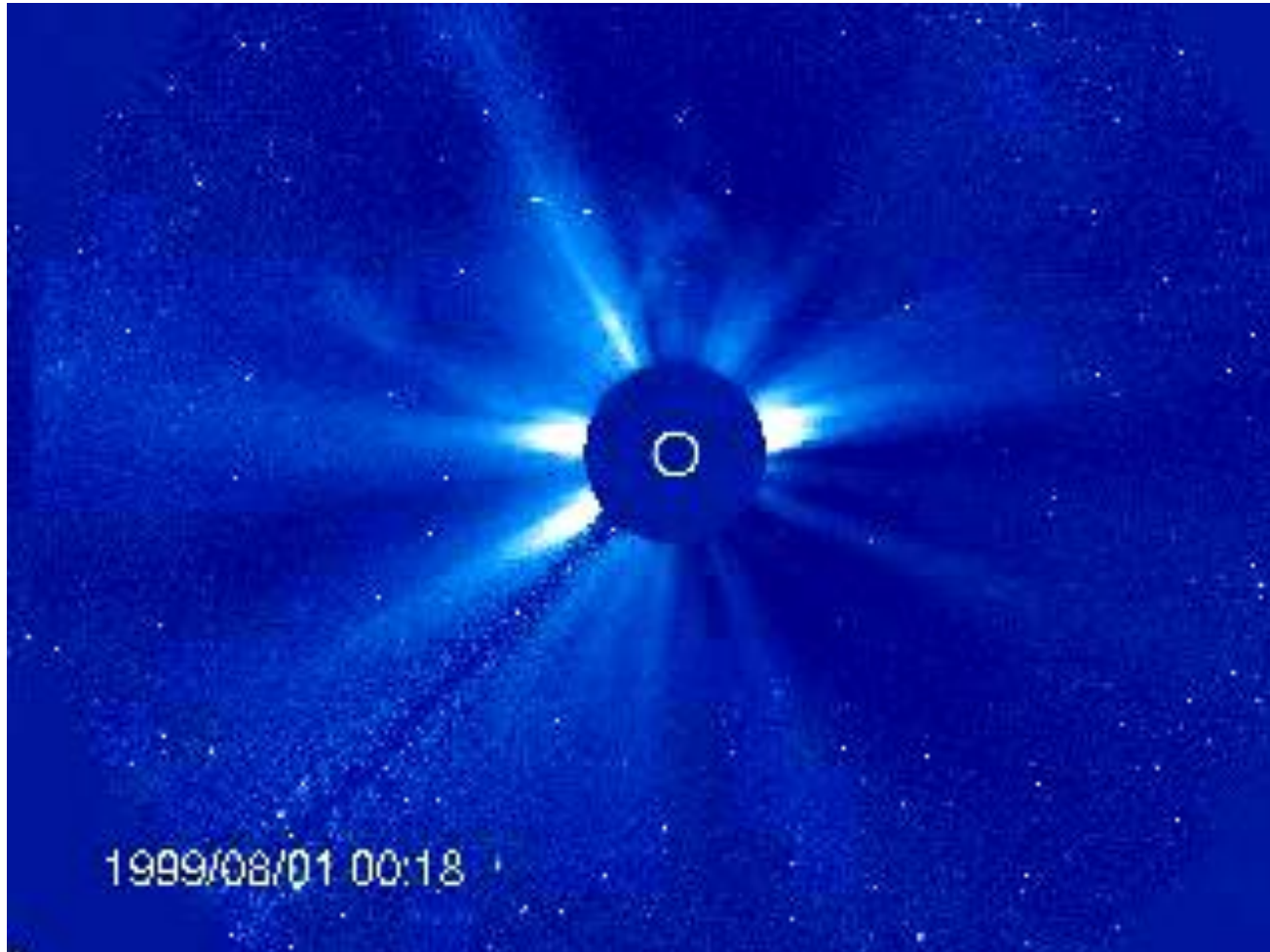
X-ray



Magnetic reconnection

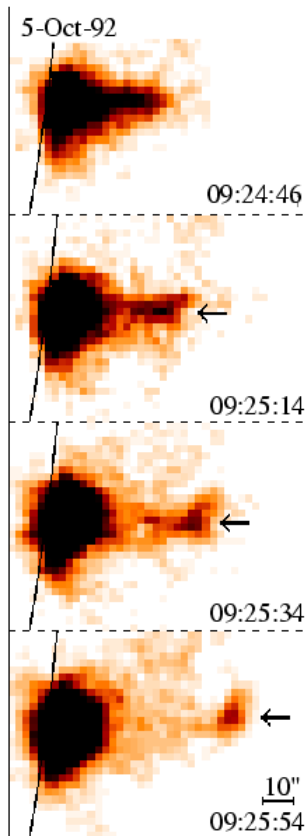


Coronal mass ejections (CME) and Solar Wind (SOHO/LASCO)

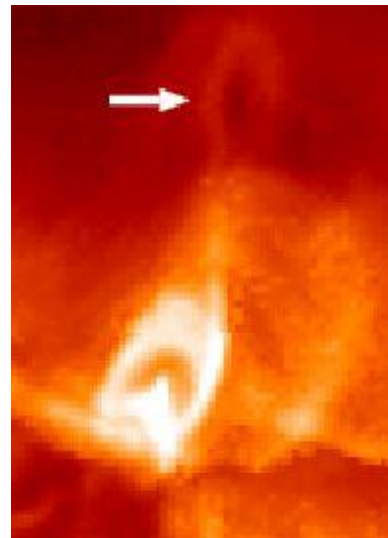


Velocity $\sim 10 - 1000$ km/s, mass $\sim 10^{15} - 10^{16}$ g

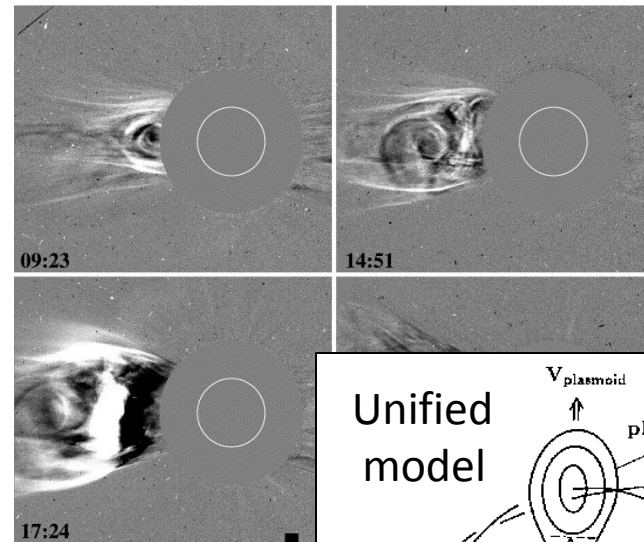
Plasmoid (flux rope) ejections



impulsive flares
~ 10^9 cm

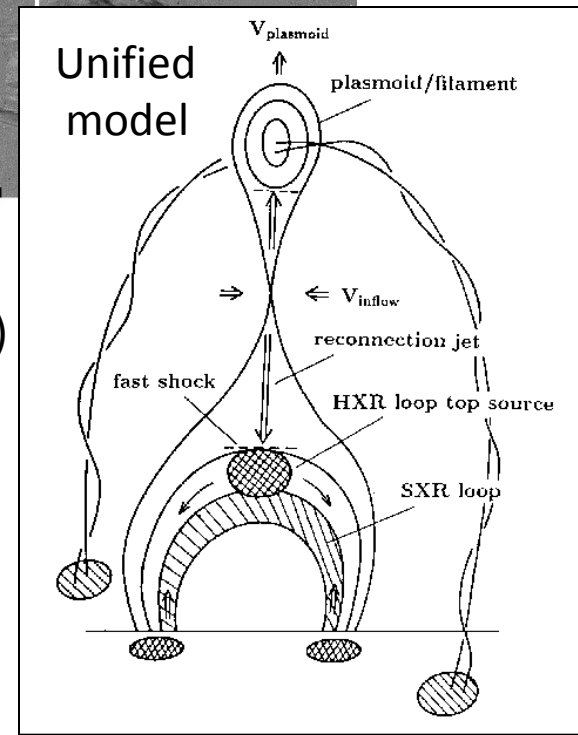


LDE(Long Duration Event) flares
~ 10^{10} cm



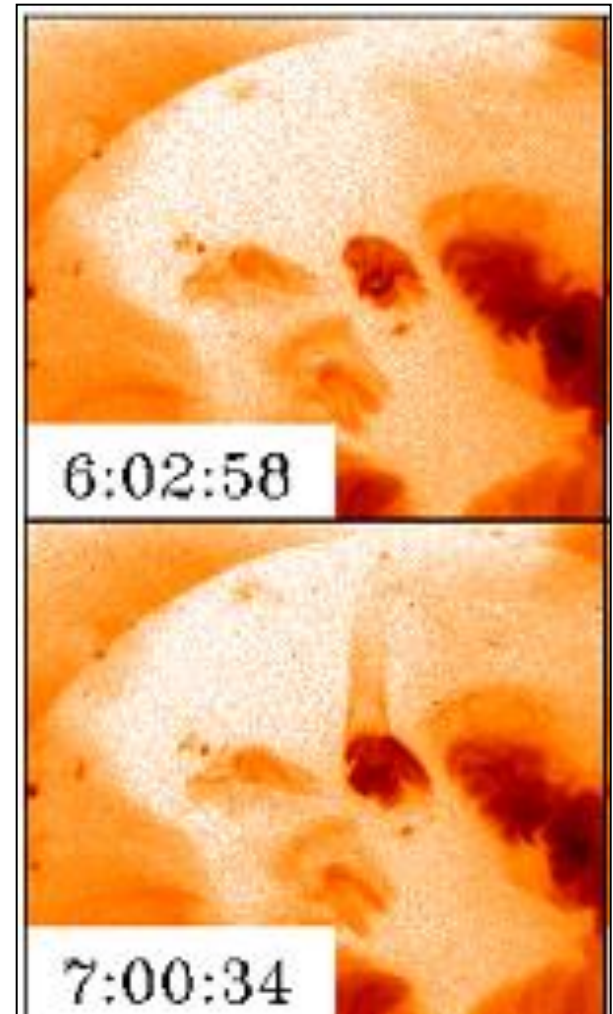
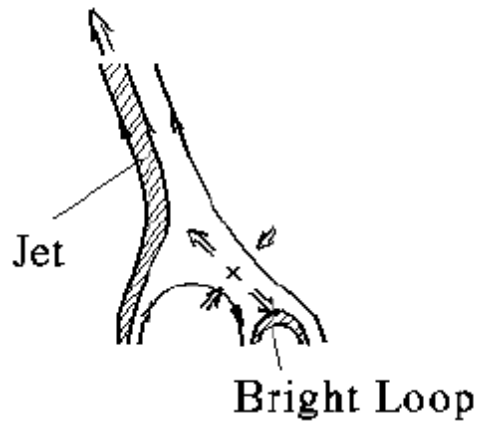
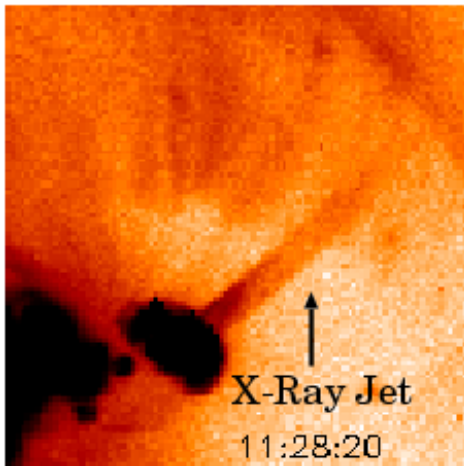
CMEs
(Giant arcades)
~ 10^{11} cm

Plasmoid-Induced-Reconnection
(Shibata 1999)



Jets from microflares

- Yohkoh/SXT discovered X-ray jet
jet
(Shibata et al. 1992,
Strong et al. 1992,
Shimojo et al. 1996)



Anemone (Shibata et al. 1994)

Summary of “flare/CME” observations with Yohkoh

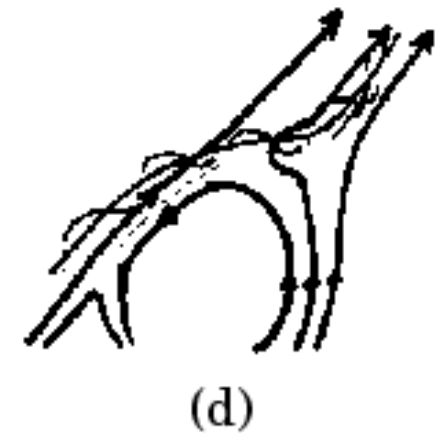
“flares”	Size (L)	Lifetime (t)	Alfvén time (t_A)	t/t_A	Mass ejection
microflares	$10^3 - 10^4$ km	100-1000sec	1-10 sec	~100	jet/surge
Impulsive flares	(1-3) x 10^4 km	10 min – 1 hr	10-30 sec	~60-100	X-ray plasmoid/ Spray
Long duration (LDE) flares	(3-10)x 10^4 km	1-10 hr	30-100 sec	~100-300	X-ray plasmoid/ prom. eruption
Giant arcades	$10^5 - 10^6$ km	10 hr – 2 days	100-1000 sec	~100-300	CME/prom. eruption

Unified model

(plasmoid-induced reconnection model)
(Shibata 1996, 1999)

(a,b): giant arcades,
LDE/impulsive flares,
CMEs

(c,d) : impulsive flares,
microflares, jets



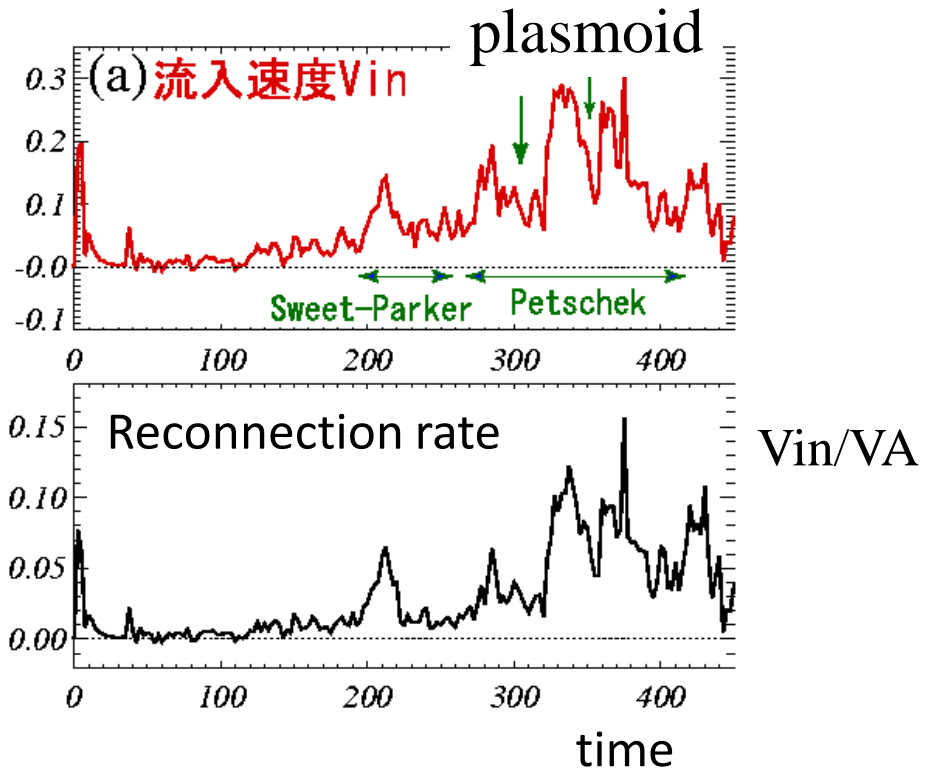
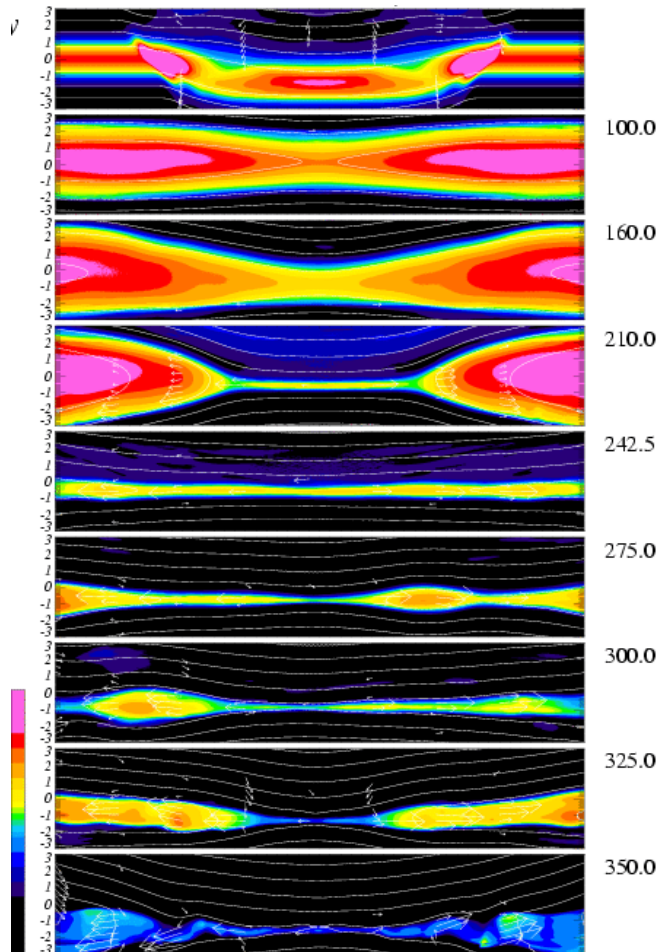
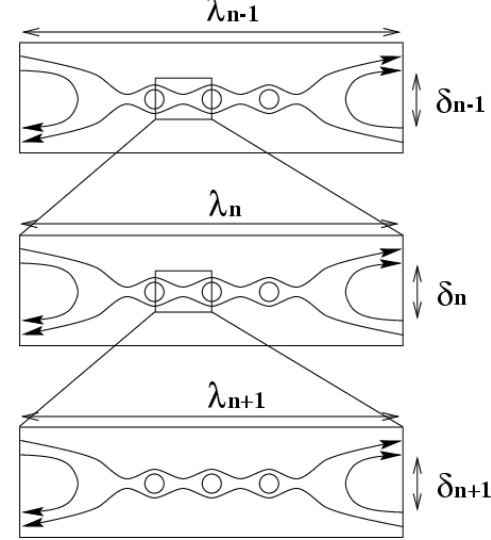
Energy release rate = $\frac{dE}{dt} \approx \frac{B^2}{4\pi} V_{in} L^2 \approx 10^{-2} \frac{B^2}{4\pi} V_A L^2$

Remaining basic questions on reconnection

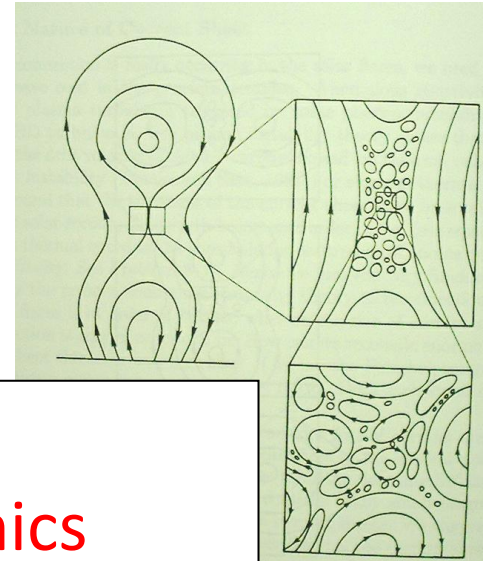
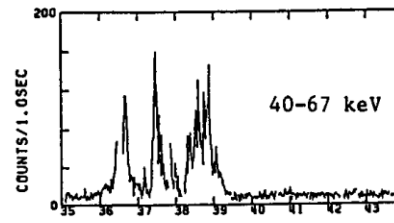
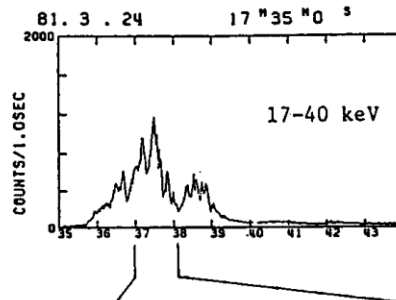
- Triggering mechanism ?
- What determines the reconnection rate ?
- What fraction of released energy goes to nonthermal particle energies ?

MHD simulations show plasmoid-induced reconnection in a fractal current sheet

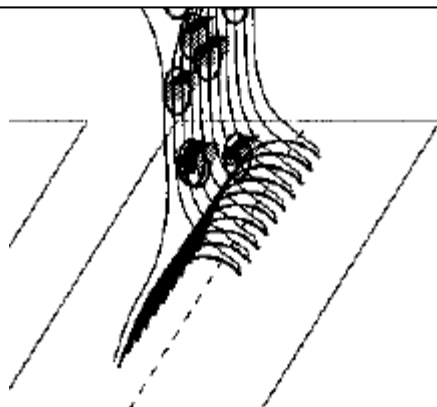
(Tanuma et al. 2001, Shibata and Tanuma 2001)



Observation of hard X-rays and microwave emissions show **fractal-like time variability**, which may be a result of fractal plasmoid ejections



This fractal structure enable to connect micro and macro scale structures and dynamics



Small-scale electric fields in magnetic X and O points

Aschwanden 2002

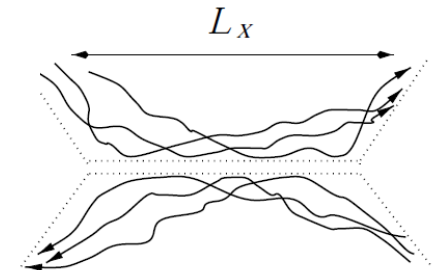
Fig.1. Time profiles of 1981 March 24 event. Two energy bands are illustrated in the upper two panels. One division of the time axis corresponds to one minute. In the bottom panel, an expanded time profile of the energy band 17 - 40 keV is shown with the time resolution of 0.125 sec.

(Ohki 1992)

Fractal current sheet

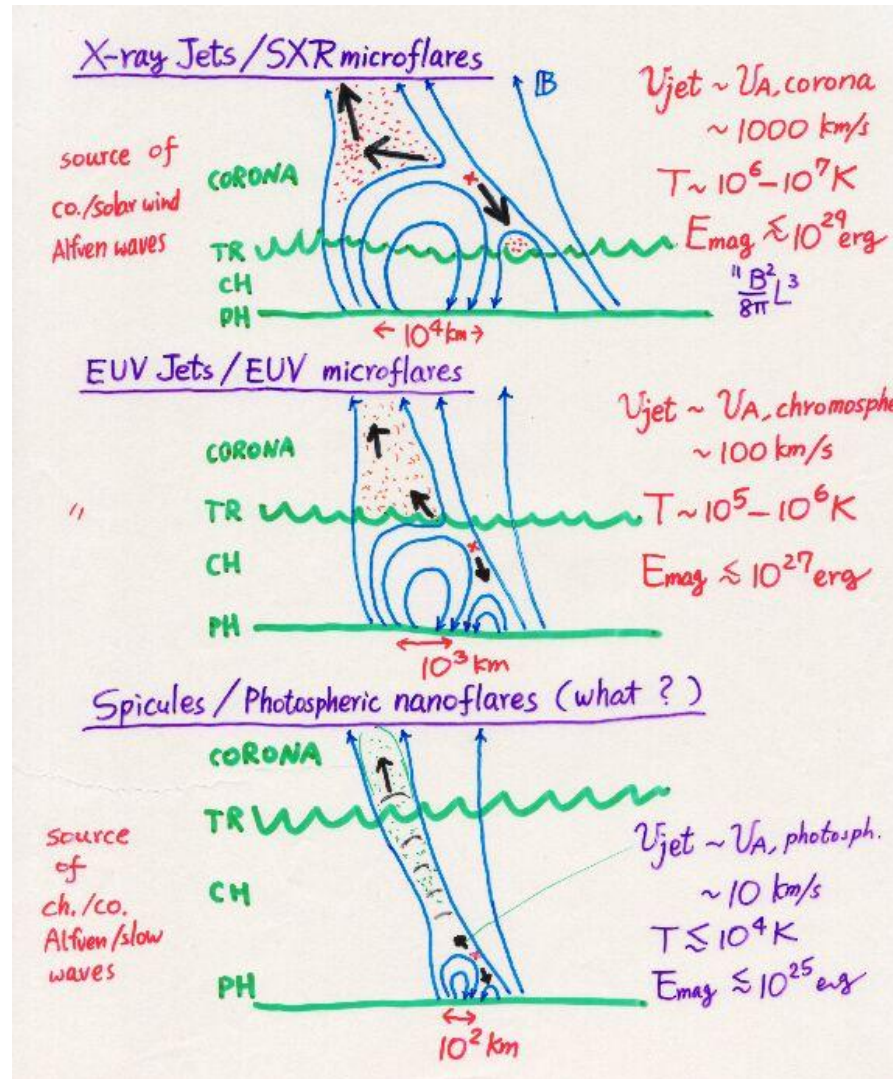
Benz and Aschwanden 1989
Zelenyi 1996, Karlicky 2004

(Tajima-Shibata 1997)



Lazarian and Vishniac 1999

Prediction of ubiquitous jets in the solar atmosphere (Shibata 1998)

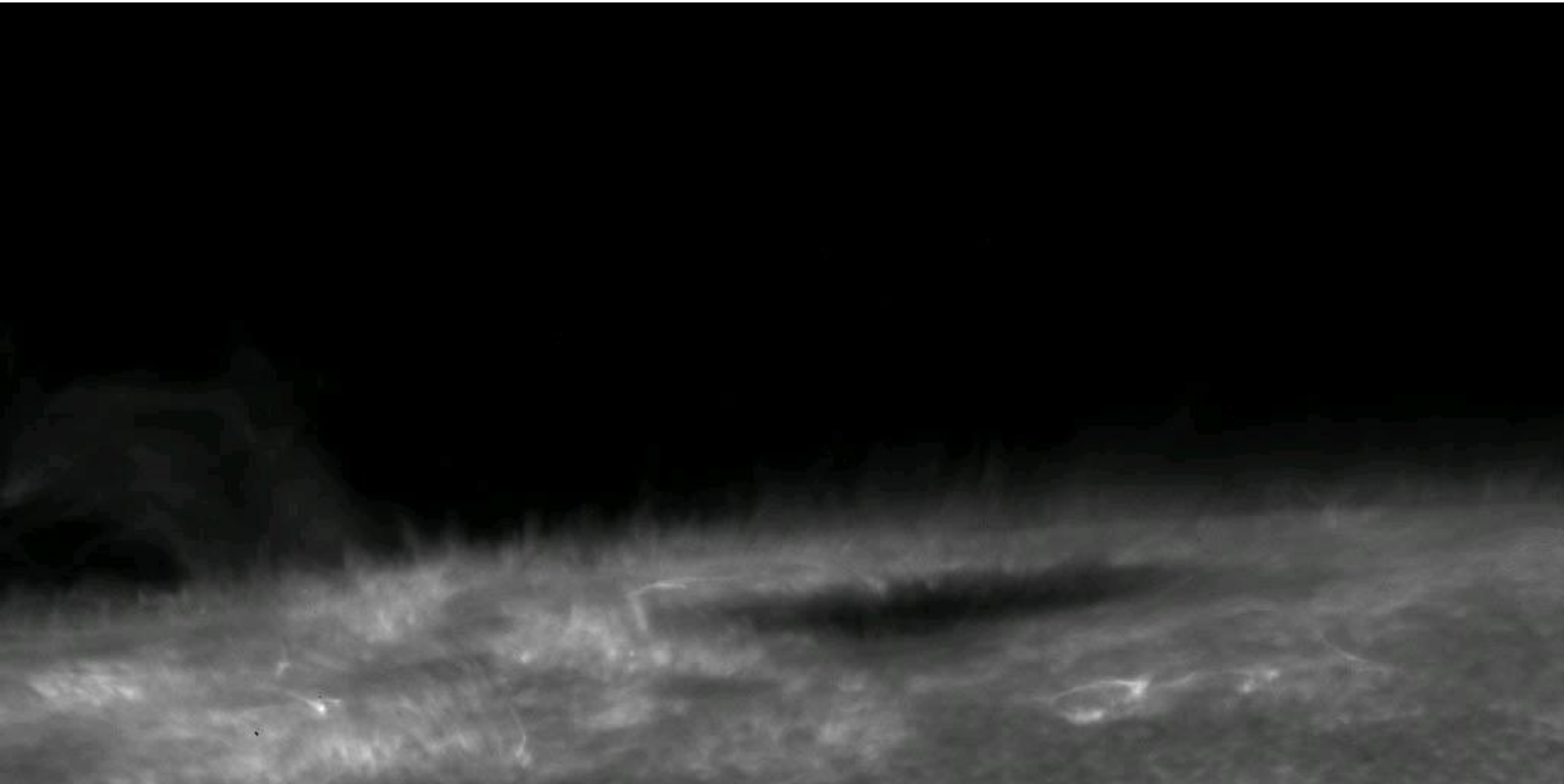


Question

- Has Hinode really observed ubiquitous cool jets ?

Answer: Yes !

Discovery of chromospheric anemone jets
(Shibata et al. 2007 Science 318, 1591)



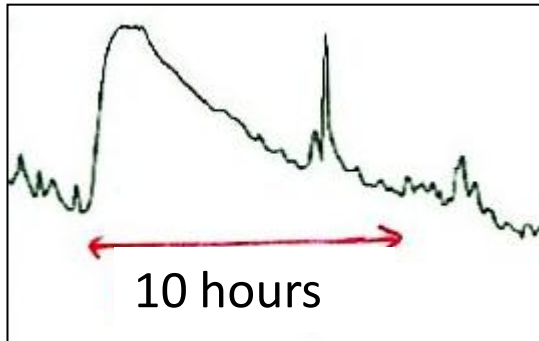
2006 Dec 17 20:00-21:00 UT Call H broad band filter images taken with Hinode/SOT

Stellar Flares

Observations of Stellar Flares

Solar Flare

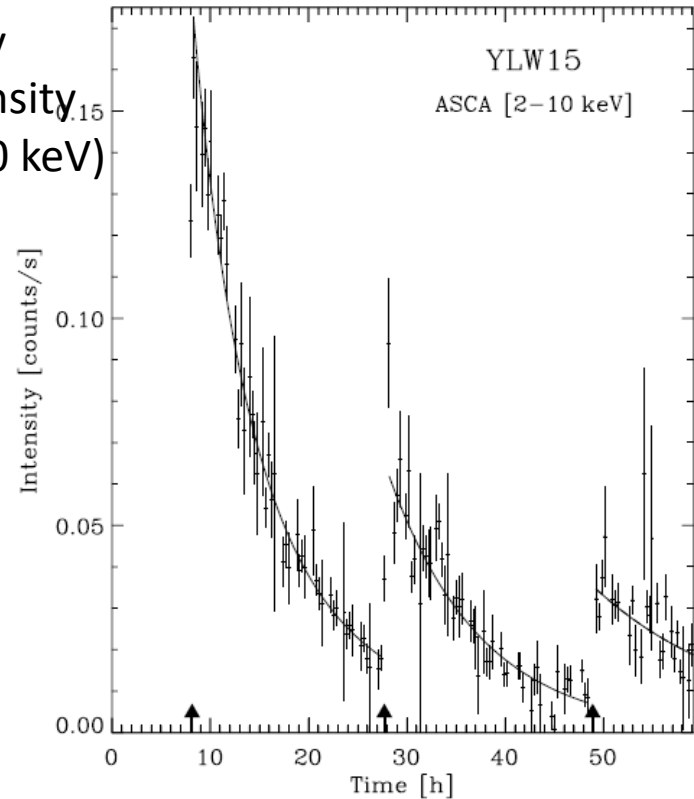
X-ray Intensity (3-24keV)



time

Protostellar Flare of YLW15

X-ray Intensity (2-10 keV)

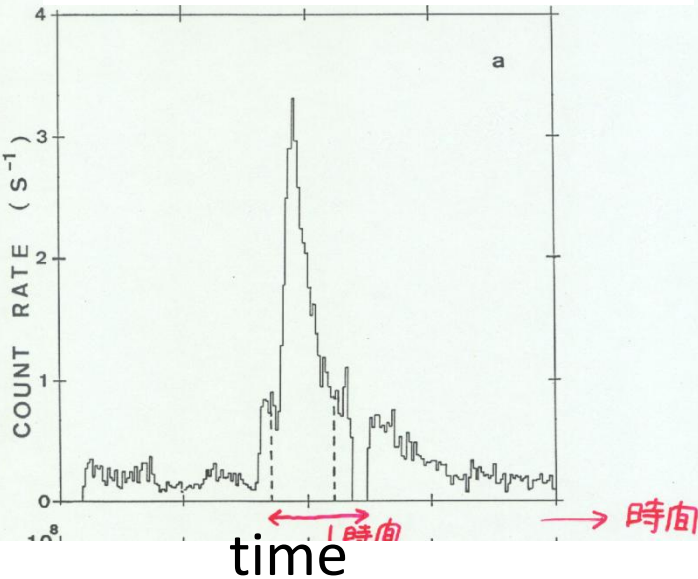


time

Stellar Flare of Prox Cen

X-ray Intensity (~ 1 keV)

X線強度



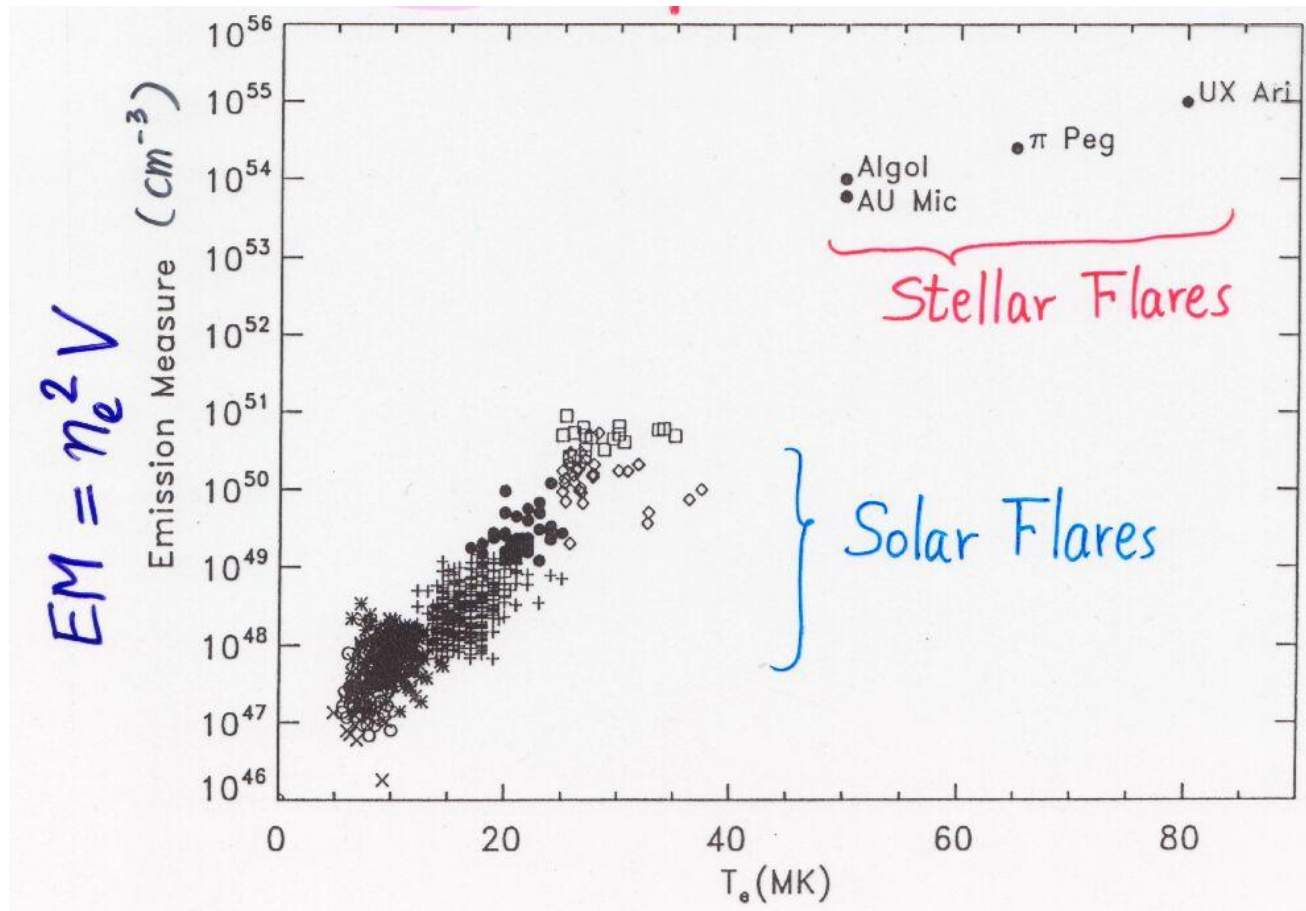
time

Can stellar and protostellar flares be explained by magnetic reconnection mechanism ?

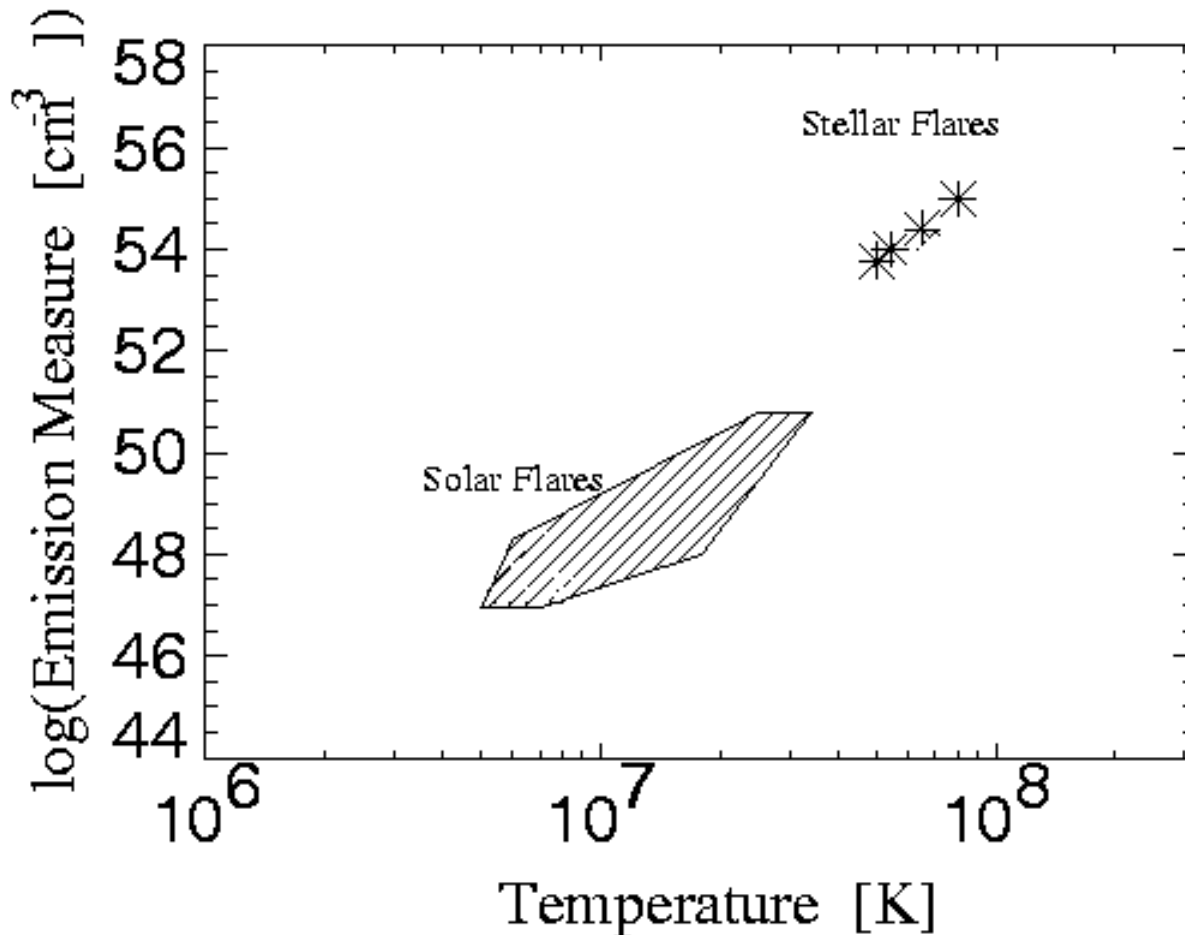
- Yes !
- Indirect evidence has been found in empirical correlation between Emission Measure ($EM = n^2 L^3$) and Temperature
(Shibata and Yokoyama 1999, 2002)

Emission Measure ($EM = n^2 V$) of Solar and Stellar Flares increases with Temperature (T)

(n : electron density, V : volume) (Feldman et al. 1995)

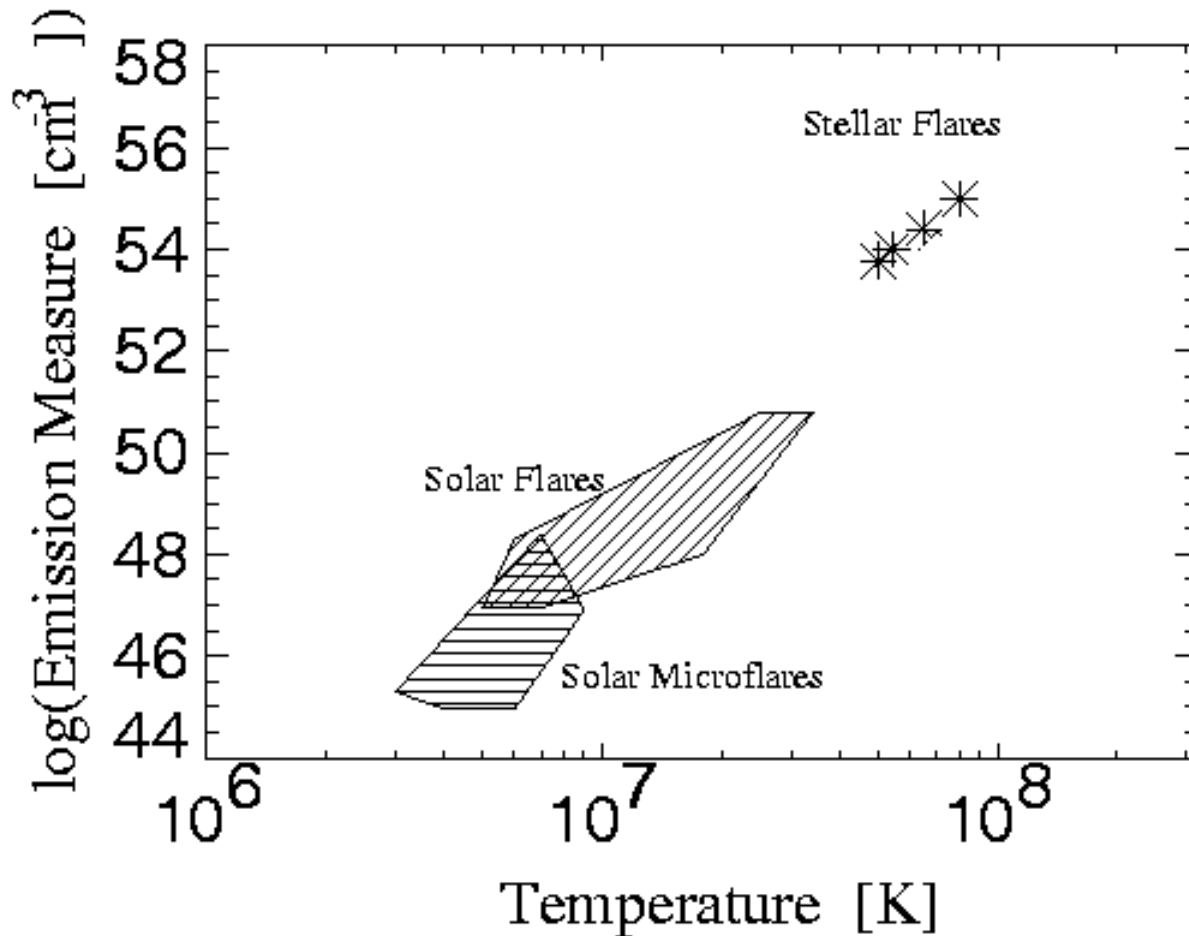


EM – T relation of Solar and Stellar Flares



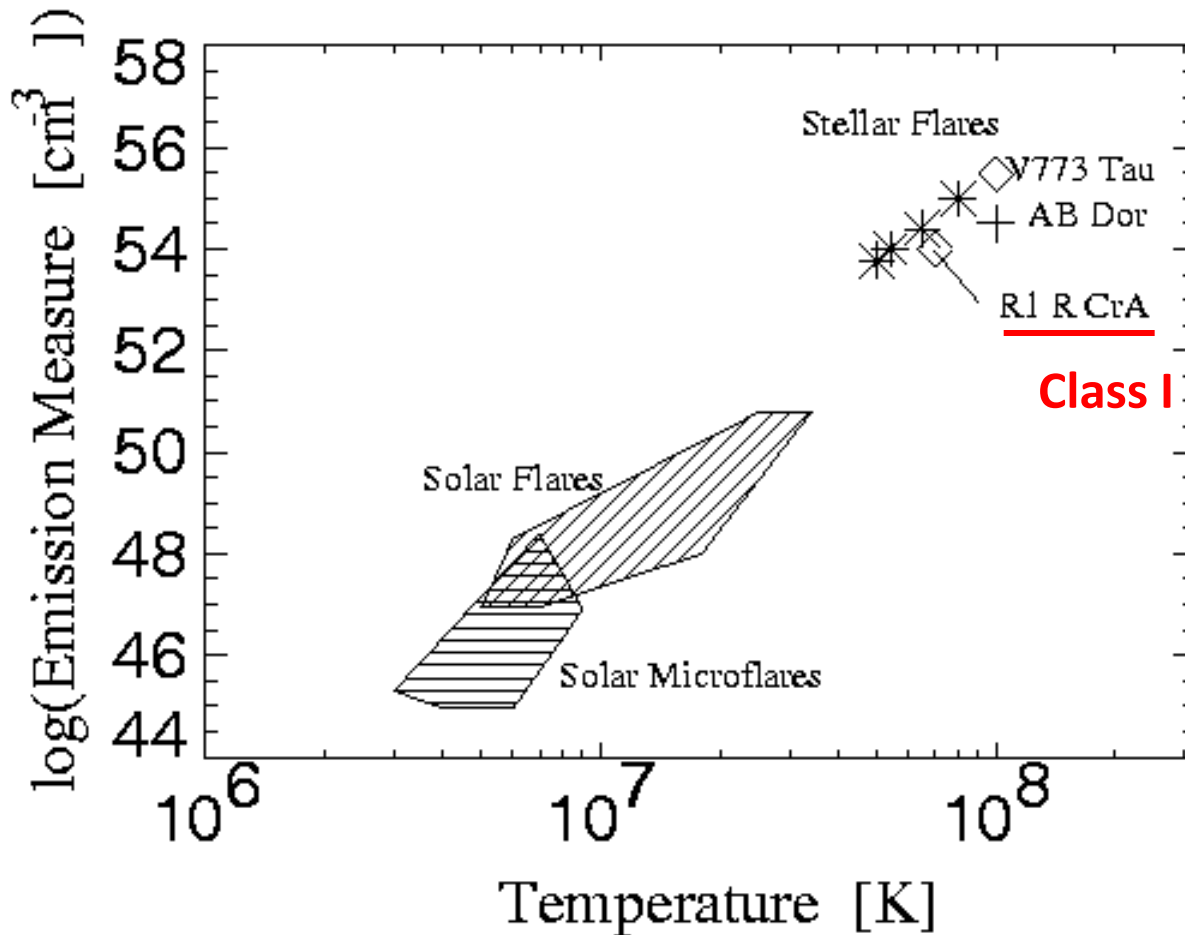
Log-log plot of Feldman et al (1995)'s figure

EM – T relation of Solar and Stellar Flares



microflare
(Shimizu 1995)

young-star and protostellar flares



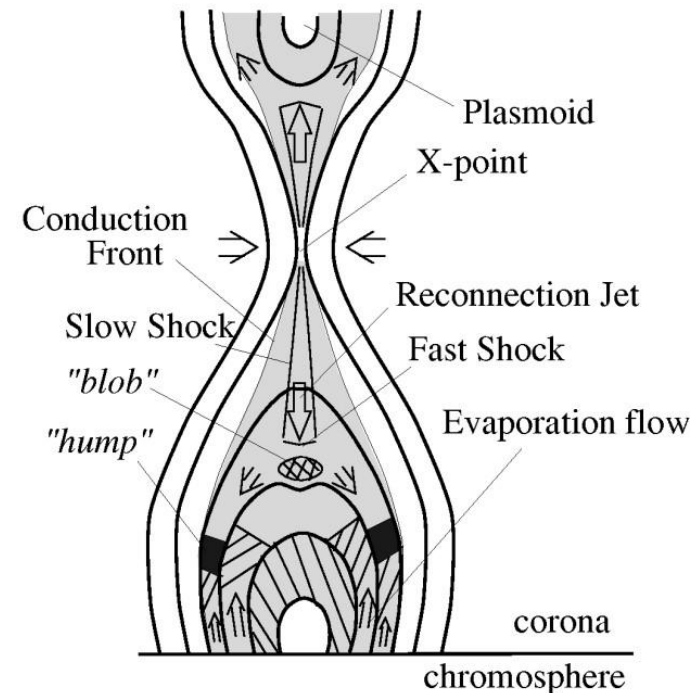
Tsuboi
(1998)
Pallavicini
(2001)
**Koyama
(1996)**
Class I protostar

Flare Temperature Scaling Law

- Reconnection heating = conduction cooling
(Yokoyama and Shibata 1998)

$$B^2 V_A / 4\pi = \kappa T^{7/2} / 2L$$

$$T \propto B^{6/7} L^{2/7}$$



Flare Emission Measure

(Shibata and Yokoyama 1999)

- Emission Measure

$$EM = n^2 L^3$$

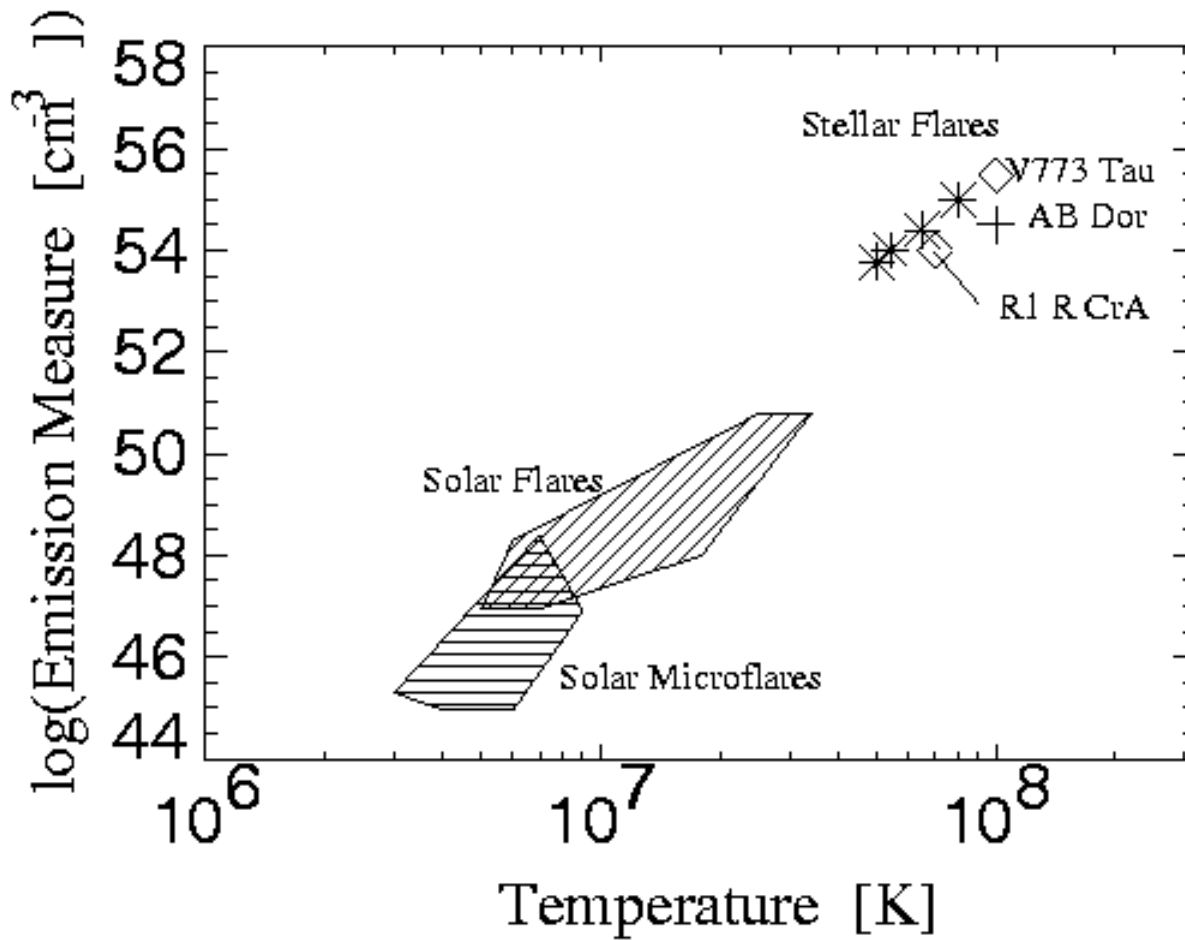
- Dynamical equilibrium (evaporated plasma must be confined in a loop)

$$2nkT = B^2 / 8\pi$$

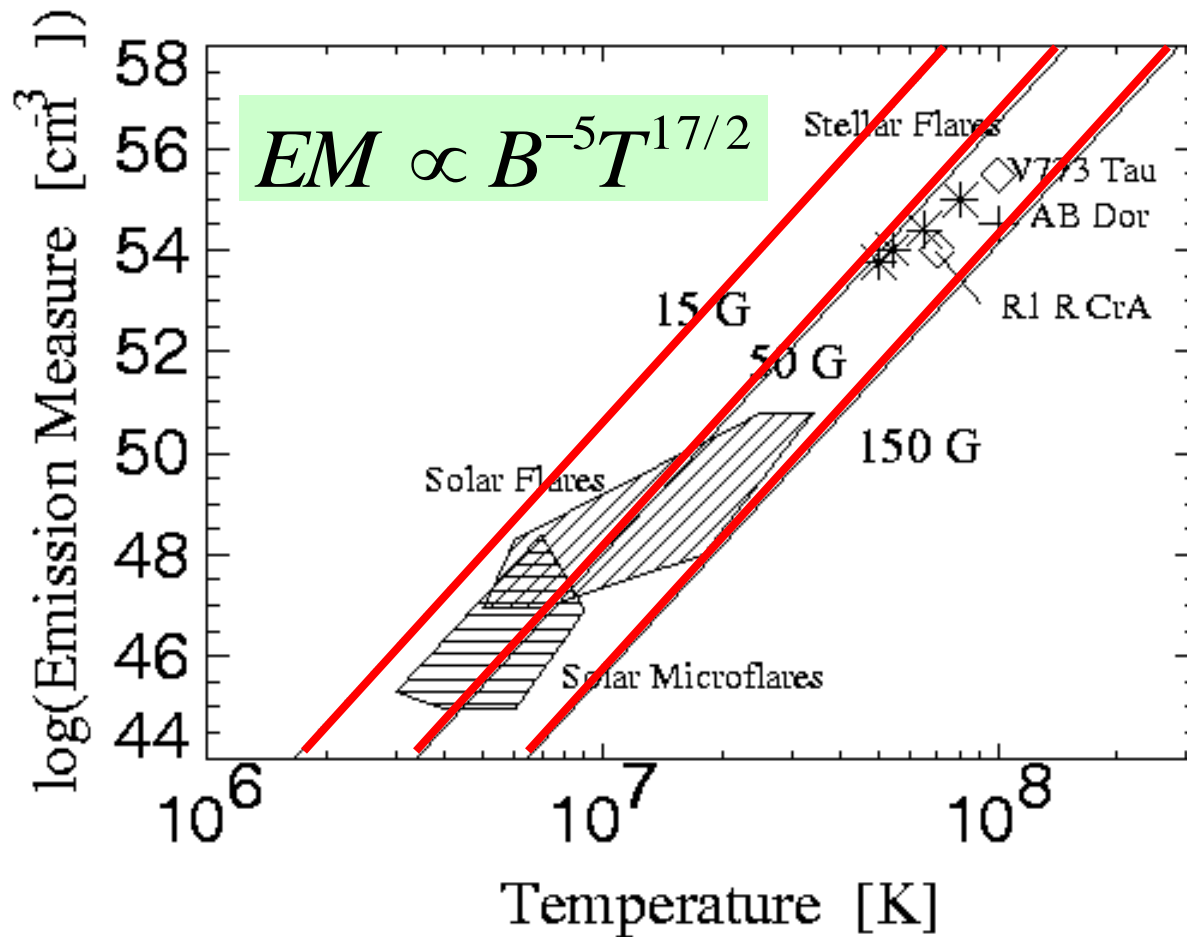
- Using Flare Temperature scaling law, we have

$$EM \propto B^{-5} T^{17/2}$$

EM-T correlation for solar/stellar flares

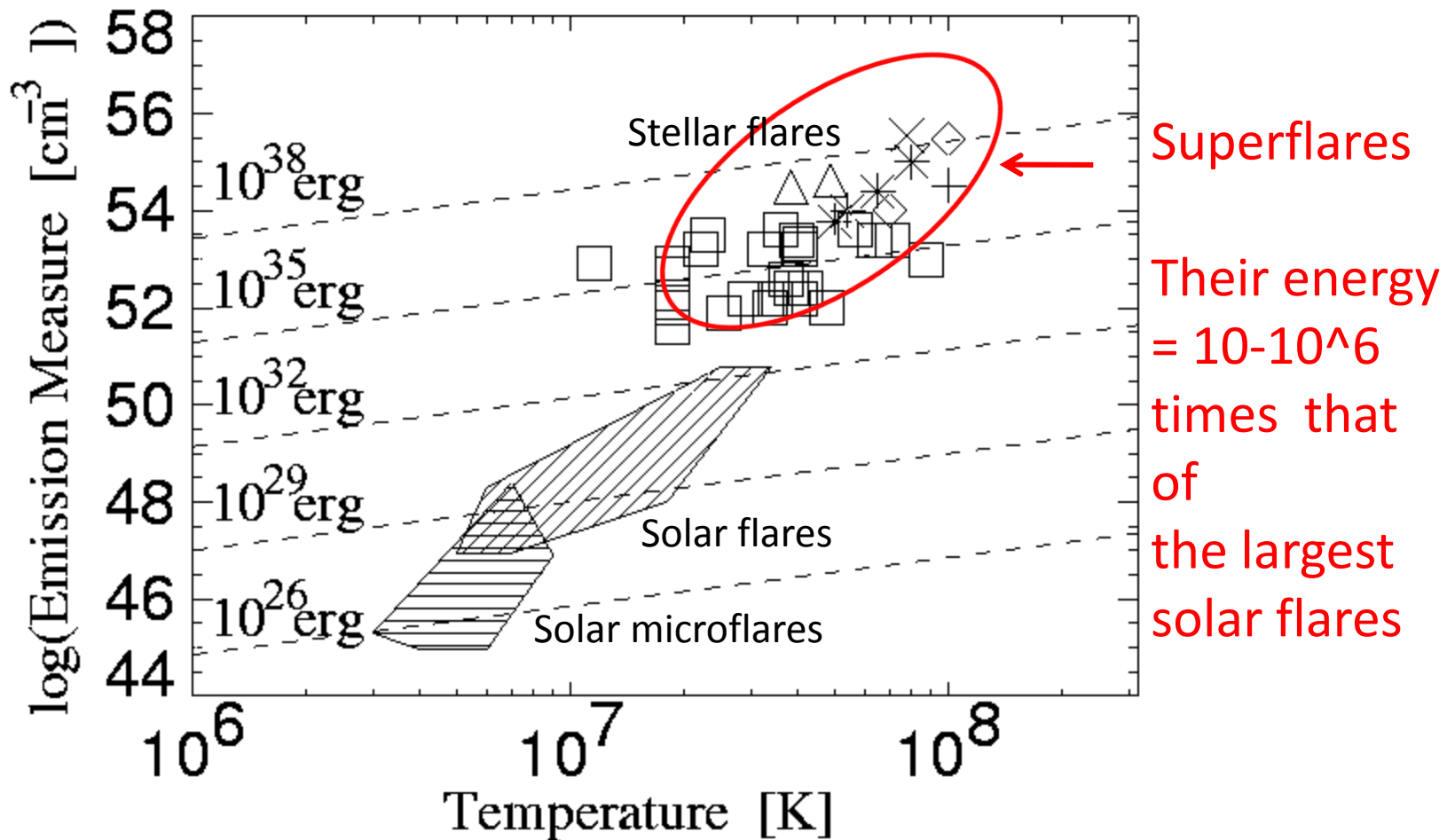


Magnetic field strength (B) = constant



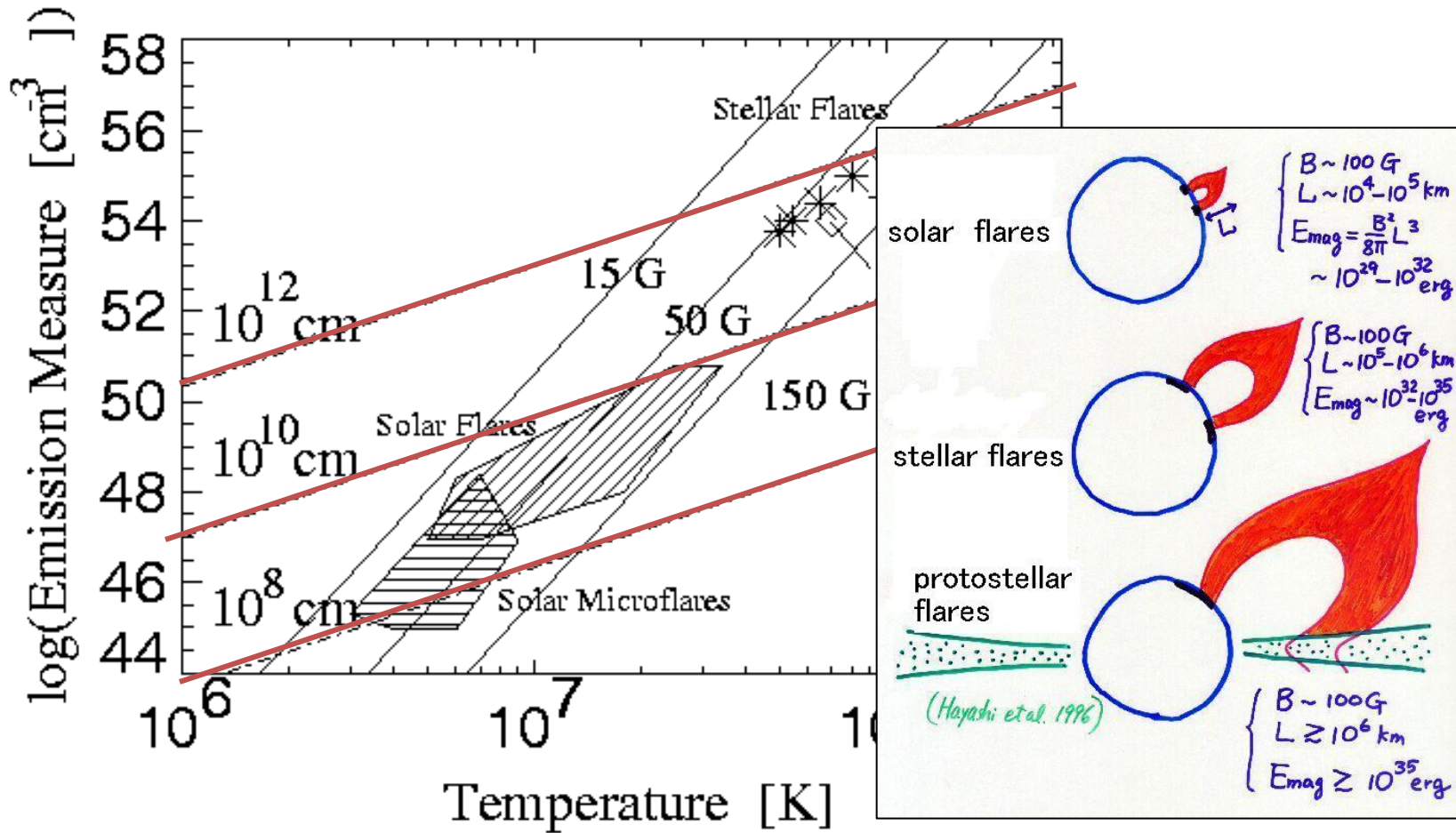
Magnetic field strengths of solar and stellar flares are comparable ~ 50 - 100 G

Total energy of stellar flares



Q: What determines the total energy of flares ?

A: It is the loop length.



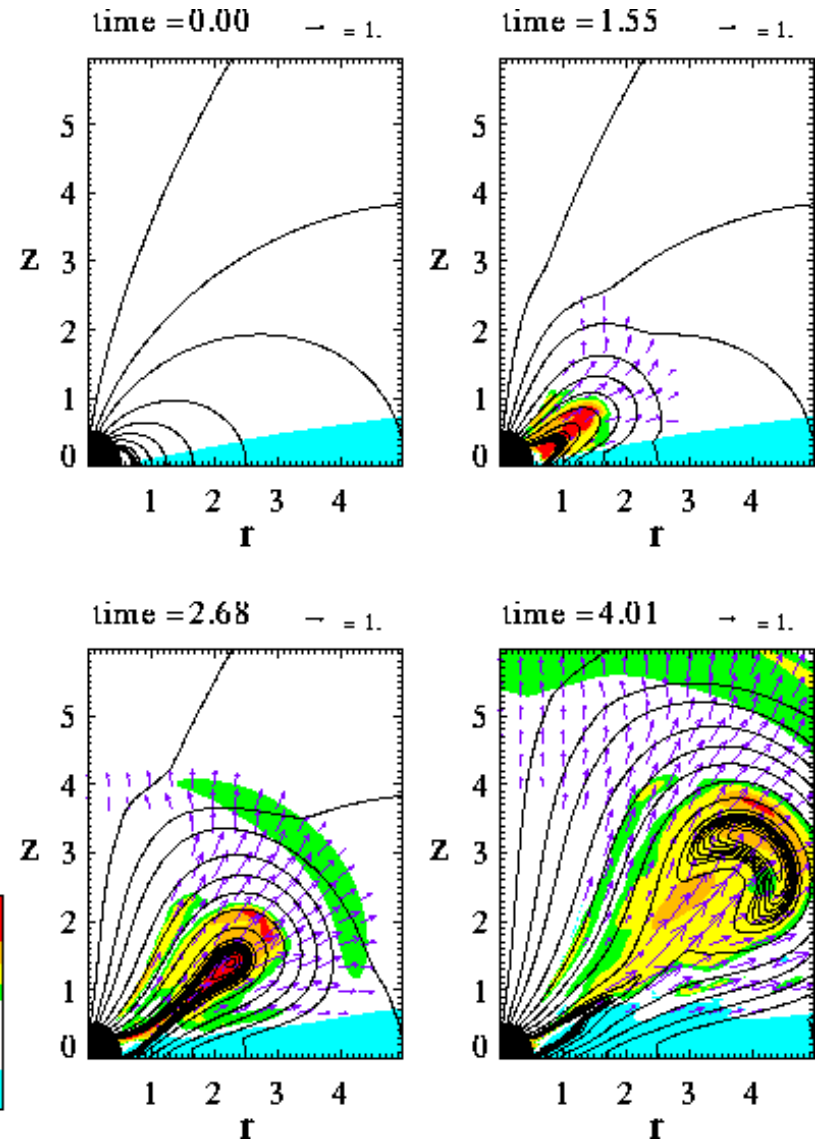
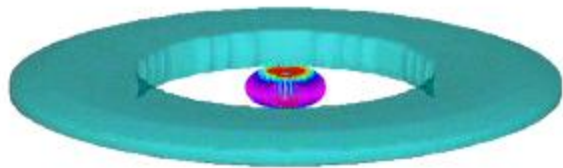
**The reason why stellar flares are hot
=> loop lengths of stellar flares are large**

Cf Isobe et al. 2003,
Kawamiti et al. 2008

reconnection model of protostellar flare and jets

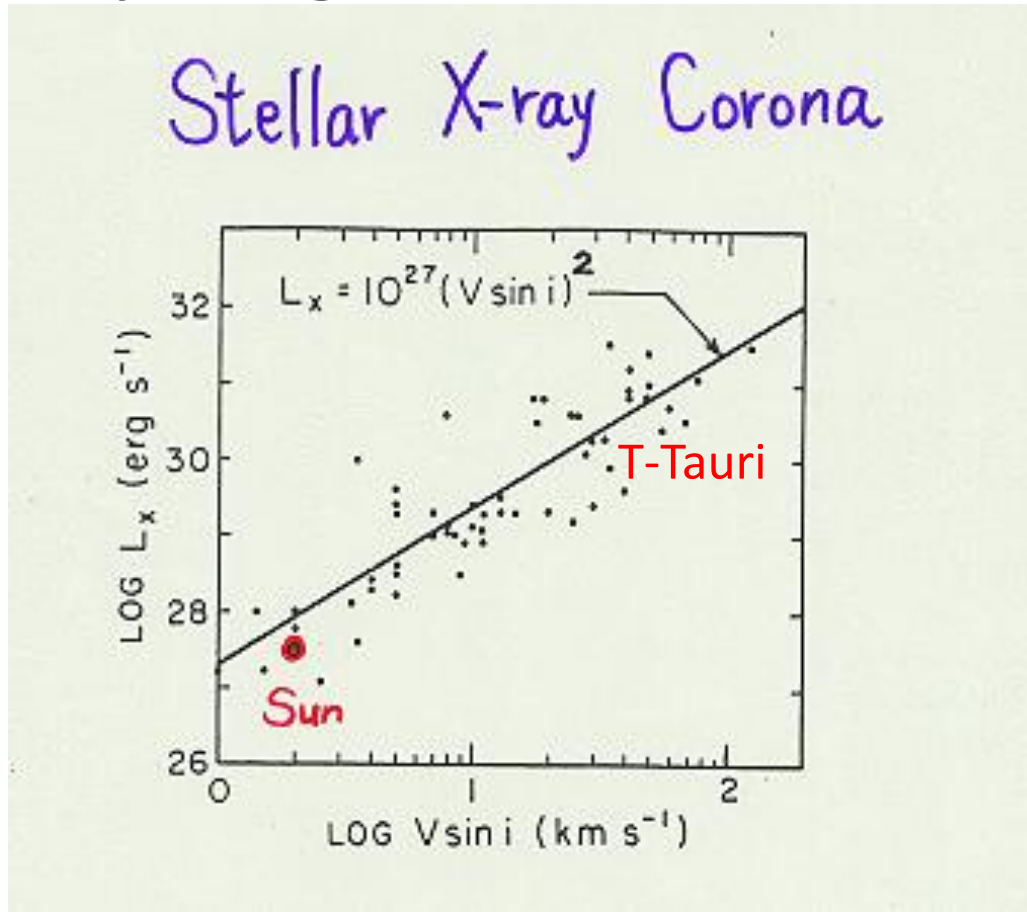
(Hayashi, Shibata,
Matsumoto 1996)

Time= 0.000 Orbits



Why young stars produce superflares ?

- Answer:
because young stars are fast rotators.



Is there a possibility that superflares would occur on our present Sun ?

Immediate Answer:

Superflares would not occur on our present Sun because the Sun is not young so is slow rotator

But amazingly, Schaefer et al. (2000) discovered 9 superflares on ordinary solar type stars with slow rotation.

Are superflares really occurring on solar type stars ?
If so, what is the occurrence frequency of superflares ?

Discovery of Superflares on Solar type stars using Kepler satellite data

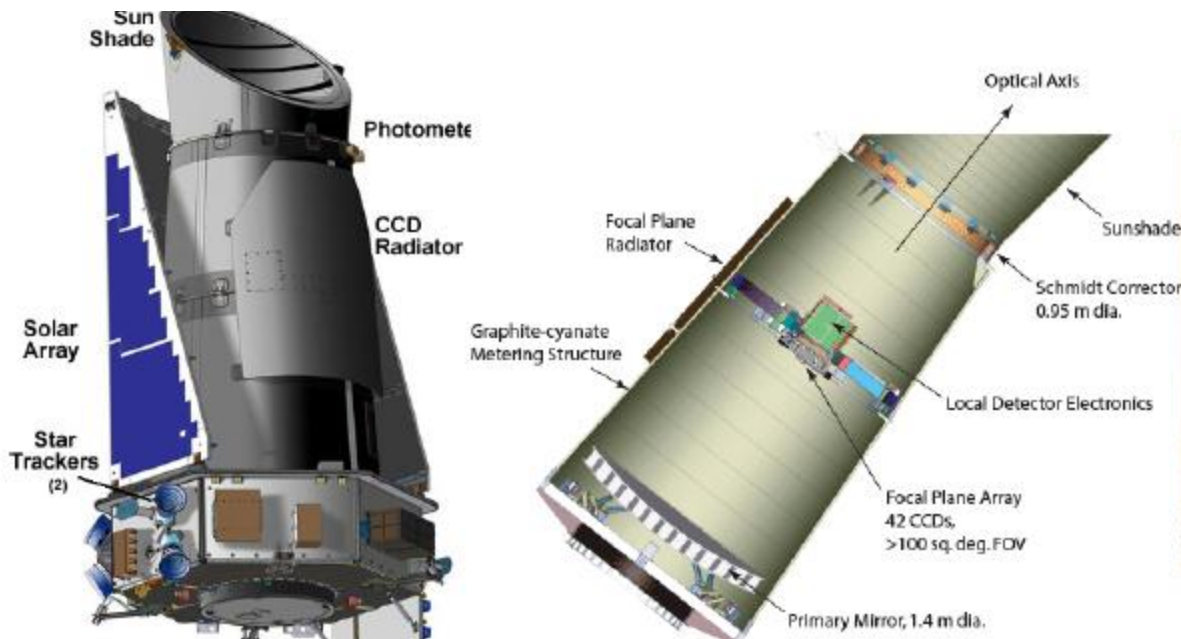
H. Maehara, et al.

(T. Shibayama, S. Notsu, Y. Notsu, T. Nagao,
S. Kusaba, S. Honda, D. Nogami, K. Shibata)

To be submitted soon (2011)

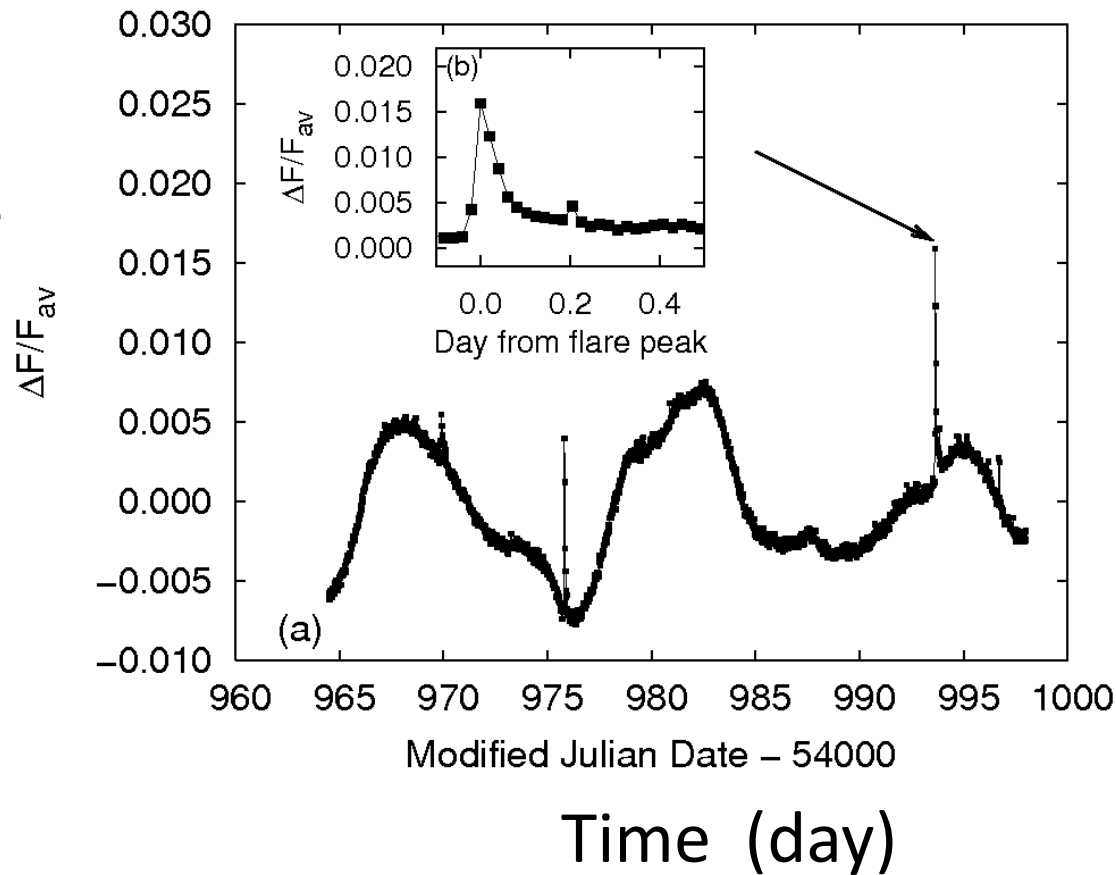
Kepler satellite

- Space mission to detect exoplanets by observing transit of exoplanets
- 42 CCD on primary focal plane of 95 cm optical telescope



typical superflare observed by Kepler

Brightness
of a star
and a flare



Total energy
 $\sim 10^{35}$ erg

Results of Superflare survey using Kepler Data

- Using the Kepler satellite data we searched for superflares on solar type stars (G type main sequence stars) and discovered **420 events**.
- More than 50 superflares are found to occur on slowly rotating stars like our Sun (its period is 25 days).
- We conclude that **superflares can occur on our Sun at present** and the occurrence frequency of the superflare whose energy is **1000** times that of the most energetic flare of the Sun is **once in 1000 years**.
- If such superflares occur on our present Sun, the Earth and our civilization would be heavily damaged.

Summary

- Recent solar observations show various evidence of reconnection in solar flares, microflares, and nanoflares, leading to a **unified model** of solar flares.
- **Plasmoid ejections** are ubiquitous in flares and flare-like events, which may play a key role to induce fast reconnection in a **fractal (turbulent) current sheet**.
- Stellar flares can also be unified with a reconnection model, if we use the emission measure – temperature diagram.
- Using Kepler data, we found that **superflares can occur on our present Sun** and the occurrence frequency of the superflare whose energy is **1000** times that of the most energetic flare of the Sun is **once in 1000 years**.